

We heuristicly discuss stationary scattering theory.

This is the typical approach of physics books & more convenient to calculate transition probabilities.

But conceptually not so clear & mathematically complicated to make sense of (but see, e.g., Reed & Simon Vol. 3).

STATIONARY SCATTERING THEORY (heuristic!)

$H_0 = -\frac{\Delta}{2}$ has formal "eigenfunctions" $\mathcal{E}_k(x) := e^{ik \cdot x}$ (plane waves).

Formally " $H_0 \mathcal{E}_k = \frac{k^2}{2} \mathcal{E}_k$ ", even though strictly speaking

$H_0 \mathcal{E}_k$ is not defined since $\mathcal{E}_k \notin L^2$!

(Of course $|e^{-iHt} \mathcal{E}_k(x)|^2 = |\mathcal{E}_k(x)|^2$ time-independently, so it is unclear how to think of this as a "flying particle". Nevertheless we think of it as a particle flying with momentum $k \in \mathbb{R}^n$; the fact that we cannot assign it a position is related to the Heisenberg uncertainty principle.)

We construct formal eigenfct. for H , that "for $t \rightarrow -\infty$ " look like plane waves,

by trying to formally (!) make sense of " $\mathcal{C} := \Omega_- \mathcal{C}_k$ ":

By intertwining we expect: $H\mathcal{C} = H\Omega_- \mathcal{C}_k = \Omega_- H_0 \mathcal{C}_k = \frac{k^2}{2} \mathcal{C}$.

Recall: for all ψ : $\Omega_- \psi = \lim_{t \rightarrow -\infty} e^{iHt} e^{-iH_0 t} \psi$.

So $\Omega_-^* \psi = \lim_{t \rightarrow -\infty} e^{iH_0 t} e^{-iHt} \psi$.

Using the abelian limit formula (see Assignment 2):

$$\begin{aligned} \Omega_-^* \psi &= \lim_{\varepsilon \downarrow 0} \varepsilon \int_{-\infty}^0 e^{\varepsilon t} e^{iH_0 t} e^{-iHt} \psi dt \\ \text{Dyson formula} &= \lim_{\varepsilon \downarrow 0} \varepsilon \int_{-\infty}^0 e^{\varepsilon t} \left[\psi + \int_0^t \frac{d}{ds} (e^{iH_0 s} e^{-iHs} \psi) ds \right] dt \\ &= \psi + \lim_{\varepsilon \downarrow 0} \varepsilon \int_{-\infty}^0 \int_0^t e^{\varepsilon t} e^{iH_0 s} (-iV) e^{-iHs} \psi ds dt \end{aligned}$$

$$\begin{aligned} &= \psi + \lim_{\varepsilon \downarrow 0} \varepsilon \int_{-\infty}^0 \underbrace{\int_s^{-\infty} e^{\varepsilon t} e^{iH_0 s} (-iV) e^{-iHs} \psi dt}_{\text{integral over } t:} ds \\ &= -\frac{1}{\varepsilon} e^{\varepsilon s} \end{aligned}$$

$$\sim \Omega_{-}^{\dagger} \psi = \psi + i \lim_{\epsilon \downarrow 0} \int_{-\infty}^0 ds e^{\epsilon s} e^{-iH_0 s} V e^{-iHs} \psi.$$

Take $\psi = \Omega_{-} \varphi_k$ (recall that Ω_{-}^{\dagger} is the left-inverse).

$$\sim \varphi_k = \psi + i \lim_{\epsilon \downarrow 0} \int_{-\infty}^0 ds e^{\epsilon s} e^{-iH_0 s} V \underbrace{e^{-iHs} \Omega_{-} \varphi_k}_{\text{intertwining} \rightarrow e^{-i\frac{k^2}{2}s} \Omega_{-} \varphi_k}$$

$$\begin{aligned} \text{resolvent identity} \left\{ \begin{aligned} &= \psi + i \lim_{\epsilon \downarrow 0} \int_{-\infty}^0 ds e^{i(H_0 - \frac{k^2}{2} - i\epsilon)s} V \psi \\ &= \psi + i \lim_{\epsilon \downarrow 0} (H_0 - \frac{k^2}{2} - i\epsilon)^{-1} V \psi. \end{aligned} \right. \end{aligned}$$

$$\sim \psi = \varphi_k - i \lim_{\epsilon \downarrow 0} (H_0 - \frac{k^2}{2} - i\epsilon)^{-1} V \psi$$

The resolvent of H_0 is given by the Yukawa potential (see Assignment 1):

$$\rightarrow \boxed{\psi(x) = e^{-ik \cdot x} - \frac{1}{2\pi} \int \frac{e^{i|k||x-y|}}{|x-y|} V(y) \psi(y) dy}$$

Lippmann-Schwinger equation

often taken as the "obvious" starting point in physics.

But: RHS still depends on ψ \rightarrow How do you solve the eqn.?

Like with the Duhamel formula, we can try to iterate to obtain a series expansion. There are many smart ways to do this (with angular momentum decomposition), see physics books. We just do a naive expansion now:

First Born approximation: We iterate by plugging in

$$\psi(x) = e^{ik \cdot x} + \mathcal{O}(V)$$

$$\Rightarrow \psi(x) = e^{ik \cdot x} - \frac{1}{2\pi} \int \frac{e^{i|k||x-y|}}{|x-y|} V(y) e^{ik \cdot y} dy + \mathcal{O}(V^2).$$

In the 1. Born approx. this is an explicit Fourier transform that can be calculated, depending on V, k and observation position x .

Physical interpretation of the Lippmann-Schwinger equation:

Think of y as small (wide supp V , \sim scale of "size of an atom") and x as large (macroscopically far away, \sim scale "meters")

Do an expansion:

$$|x-y| = |x| \left| \frac{x}{|x|} - \frac{y}{|x|} \right|, \quad \text{Taylor expansion} \approx \frac{|y|}{|x|}:$$

$$\frac{e^{i|k||x-y|}}{|x-y|} = \frac{e^{i|k||x|}}{|x|} e^{-ik' \cdot y} + \mathcal{O}\left(\frac{|y|}{|x|}\right)$$

where $k' := |k| \frac{x}{|x|} \Rightarrow$ the "wave vector" in the direction of the observation.

Then the Lippmann-Schwinger equation takes the following form:

$$\psi(x) \approx e^{ik \cdot x} + \underbrace{f(k, k')}_{\text{scattering amplitude}} \underbrace{\frac{e^{i|k||x|}}{|x|}}_{\text{outgoing spherical wave}}$$

\uparrow
 incoming plane wave

Interpretation: If we shoot a particle at the target potential V with energy $|k|^2/2$ from direction k , then $|f(k, k')|^2$ is the probability to detect it later far away in the direction of k' .
 (The energy is of course conserved, $|k'|^2/2 = |k|^2/2$.)

VII LOCALIZATION OF BOUND STATES

We have seen: in asymptotically complete systems, the Hilbert space can be decomposed into a direct sum of bound states and scattering states, $\mathcal{H} = \mathcal{H}_B \oplus \text{ran } \Omega_+$, where $\mathcal{H}_B = \text{closure of } \{ \psi \in D(H) : H\psi = E\psi \}$. eigenvectors

Scattering states behave like free states: particle flies away and disperses.

:f $H = H_0 + V$, $\sigma(H) \subset \mathbb{R}$, so $E \in \mathbb{R}$.

Eigenvectors (bound states) have trivial time evolution:

$$H\psi = E\psi \Rightarrow e^{-iHt}\psi = e^{-iEt}\psi.$$

The probability to find a particle near position $x \in \mathbb{R}^n$:

$$|(e^{-iHt}\psi)(x)|^2 = |\psi(x)|^2 \quad \text{indep. of time.}$$

Goal: Show that $|\psi(x)|^2 \simeq e^{-C|x|}$ for large $|x|$.

PROP.: (IMS localization formula)

Let $V: \mathbb{R}^n \rightarrow \mathbb{R}$ measurable and such that $H = -\frac{\Delta}{2} + V$ is closed with domain $H^2(\mathbb{R}^n)$.

Let $f \in C^\infty(\mathbb{R}^n)$, real-valued and $\partial^\alpha f \in L^\infty(\mathbb{R}^n) \forall |\alpha| \leq 2$.

Then on $H^2(\mathbb{R}^n)$ we have

$$2fHf = f^2 H + Hf^2 + |\nabla f|^2.$$

PROOF: We only show $2fHf\varphi = (f^2H + Hf^2 + |df|^2)\varphi$ $\forall \varphi \in C_0^\infty(\mathbb{R}^n)$.

The identity for $\varphi \in H^2(\mathbb{R}^n)$ follows by an approximation argument in the graph norm of H .

(A priori it is not clear that $-\Delta = \mathcal{F}^{-1}p^2\mathcal{F}$ satisfies a product rule; only when applied to C^2 -functions, where it agrees with the classical derivatives.)

Straightforward calculation:

$$2f(-\frac{\Delta}{2} + V)f\varphi = f^2(-\frac{\Delta}{2} + V)\varphi + (-\frac{\Delta}{2} + V)f^2\varphi + |df|^2\varphi.$$

(because $f\varphi \in D(V) \cap D(-\frac{\Delta}{2})$; and $fVf\varphi$ can be written out pointwise almost everywhere.)

Use product rule for Laplacian:

$$\begin{aligned} \text{LHS: } f\Delta(f\varphi) &= f\Delta(\varphi + f\varphi) \\ &= f(\Delta\varphi) + f(\Delta f)\varphi + f(\nabla f)(\nabla\varphi) + f(\nabla f)(\nabla\varphi) + f^2(\Delta\varphi). \end{aligned}$$

$$\begin{aligned} \text{RHS: } \frac{1}{2}f^2(\Delta\varphi) + \frac{1}{2}\Delta(f^2\varphi) - |df|^2\varphi \\ = \dots = \text{LHS.} \end{aligned}$$

↪ Check it!



THM.: (Exp. localization of eigenvectors, Agmon)

Let $V \in L^2_{loc}(\mathbb{R}^n)$ and $n \geq 3$, or $V \in L^\infty(\mathbb{R}^n)$.

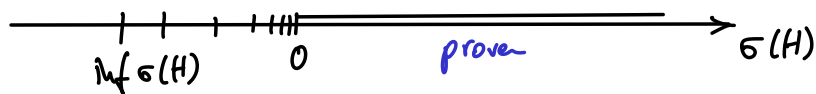
Let V real-valued and $V(x) \rightarrow 0$ ($|x| \rightarrow \infty$).

Let $H = -\frac{\Delta}{2} + V$.

If $\psi \in H^2(\mathbb{R}^n)$ and $H\psi = E\psi$, then

$$e^{\beta|x|} \psi \in L^2(\mathbb{R}^n) \quad \forall \beta > 0 \text{ with } E + \frac{\beta^2}{2} < 0.$$

RMK: $\sigma(-\frac{\Delta}{2}) \subset \sigma(H)$ if Ω_\pm exist. $\sigma(-\frac{\Delta}{2}) = [0, \infty)$.



eigenvalues, possibly accumulating at zero.
(not proven so far)

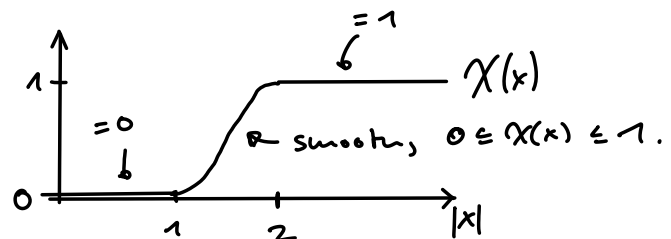
This means for the decay of eigenvectors, $H\psi = E\psi$:

$$\psi(x) \sim e^{-\beta|x|} \text{ for } |x| \rightarrow \infty,$$

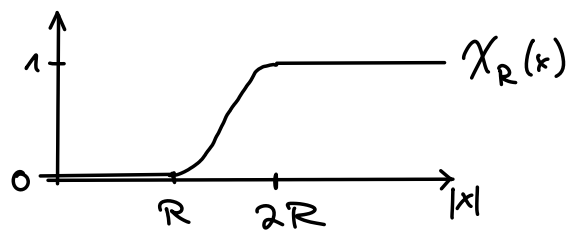
but rate of decay β smaller and smaller
for E closer and closer to zero.

PROOF (of exp. localization):

Let $\chi \in C^\infty(\mathbb{R}^n)$:



Let $\chi_R(x) := \chi(\frac{x}{R})$:
(for $R > 0$)



$$\text{let } f(x) := \frac{\beta|x|}{1 + \varepsilon|x|}$$

(for $\varepsilon > 0$ this is a regularization of $\beta|x|$).

Properties (check!):

$$|f| \leq \frac{\beta}{z}, \text{ and for } x \neq 0: |\nabla f(x)| \leq \beta.$$

(forgetting +1 in denominator)

(growth \searrow slower than of $\beta|x|$,
and the growth of $\beta|x|$ is
proportional to β)

Let $G := \chi_R e^f$. Then $G \in C^\infty(\mathbb{R}^n) \cap L^\infty(\mathbb{R}^n)$,

$$(*) \quad \nabla G = (\nabla \chi_R) e^f + \chi_R e^f (\nabla f) = (\nabla \chi_R) e^f + G (\nabla f) \in L^\infty(\mathbb{R}^n),$$

and $\partial_i \partial_j G \in L^\infty(\mathbb{R}^n) \quad \forall i, j \in \{1, \dots, n\}$.

$$\text{Now: } \langle G e, (H-E) G e \rangle \stackrel{\text{MS}}{=} \langle e, \left(\frac{1}{2} [G^2 H + H G^2 + |\nabla G|^2] - E \right) e \rangle$$

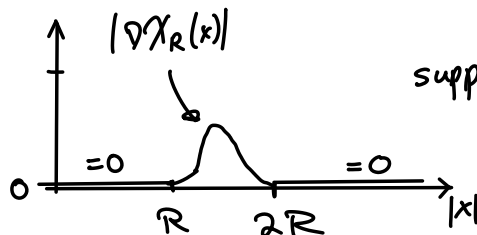
$$\stackrel{H^2 = H, H e = E e}{=} \langle e, \frac{1}{2} |\nabla G|^2 e \rangle$$

$$\stackrel{(*)}{=} \frac{1}{2} \langle e, (|\nabla \chi_R|^2 e^{2f} + 2 \nabla \chi_R \cdot \nabla f e^f G + G^2 |\nabla f|^2) e \rangle$$

← to the other side.

$$\Rightarrow \langle G e, (H-E - \frac{|\nabla f|^2}{2}) G e \rangle \leq \sup_{x \in \mathbb{R}^n} (|\nabla \chi_R|^2 e^{2\beta|x|} + 2 |\nabla \chi_R| \beta e^{2\beta|x|}) \|e\|_{L^2}^2.$$

Notice:



$$\text{supp } |\nabla \chi_R(x)| \subset \{x \in \mathbb{R}^n : R \leq |x| \leq 2R\}.$$

$$\ln \text{supp } |\nabla \chi_R(x)|: e^{2\beta|x|} \leq e^{4\beta R} < \infty.$$

$$\Rightarrow \langle G e, (H-E - \frac{|\nabla f|^2}{2}) G e \rangle \leq C_R \|e\|_{L^2}^2, \quad C_R < \infty. \quad (i)$$

On the other hand:

$$\langle G\psi, (H - E - \frac{10f^2}{2})G\psi \rangle \geq \langle G\psi, (V - E - \frac{10f^2}{2})G\psi \rangle$$

$$\langle \psi, (-\Delta)\psi \rangle = \int dp |\hat{\psi}(p)|^2 p^2 \geq 0.$$

$$|0f| \leq \delta \Rightarrow \langle G\psi, (V - E - \frac{\delta^2}{2})G\psi \rangle$$

$$\text{supp } G \subset \mathbb{R}^n \setminus B_R(0) \Rightarrow \left(\inf_{|x| \geq R} V(x) - E - \frac{\delta^2}{2} \right) \|G\psi\|_{L^2}^2$$

Recall: $E < 0$.
 > 0 for some $\delta > 0$
 because $V(x) \rightarrow 0$ ($|x| \rightarrow \infty$).

$$\geq \delta_R \|G\psi\|_{L^2}^2. \quad (ii)$$

$$\text{Combining (i) \& (ii)} \Rightarrow \|G\psi\|_{L^2}^2 \leq \frac{C_R}{\delta_R} \|\psi\|_{L^2}^2.$$

estimate uniform in $\varepsilon > 0!$

$$\Rightarrow \int_{|x| \geq 2R} e^{2\beta|x|} |\psi(x)|^2 dx = \int_{|x| \geq 2R} \lim_{\varepsilon \downarrow 0} e^{2f(x)} |\psi(x)|^2 dx$$

Lebesgue
monotone
convergence

$$= \lim_{\varepsilon \downarrow 0} \int_{|x| \geq 2R} e^{2f(x)} |\psi(x)|^2 dx$$

$$\leq \lim_{\varepsilon \downarrow 0} \int \underbrace{X_R^2(x)}_{= G(x)^2} e^{2f(x)} |\psi(x)|^2 dx$$

charact. fun. of the
ball $B_{2R}(0)$ is
bounded by $X_R(x)^2$
(look at the pictures above).

$$\leq \lim_{\varepsilon \downarrow 0} \frac{C_R}{\delta_R} \|\psi\|^2 = \frac{C_R}{\delta_R} \|\psi\|^2.$$

$$\begin{aligned} \int_{|x| \leq 2R} e^{2\beta|x|} |\varphi(x)|^2 dx &\leq \int_{|x| \leq 2R} e^{4R\beta} |\varphi(x)|^2 dx \\ &\leq e^{4R\beta} \|\varphi\|_{L^2}^2 < \infty. \end{aligned}$$

$$\Rightarrow e^{\beta|x|} \varphi \in L^2(\mathbb{R}^n).$$



RMK: One can strengthen the result to
 $|\varphi(x)| \leq C_\beta e^{-\beta|x|}$ almost everywhere.